MODELING AND DESIGN OF SOLAR ENERGY SYSTEMS INCLUDING SOLAR ECONOMICS

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Keywords: F-chart method, liquid solar energy systems, air solar energy systems, service water systems, thermosiphon systems, modeling, simulations, simple models, TRNSYS

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Summary

The design of a solar energy system is a complex problem. As such it includes both predictable and unpredictable parameters. The former are related to performance characteristics of collectors and other components and the latter mainly concern weather data such as solar radiation, ambient temperature, wind speed and direction, and other

less important environmental parameters. In this chapter, a simple design method and a simple modeling technique are presented as well as an overview of TRNSYS which is the simulation program suitable for modeling active solar heating systems. The design method presented is the *f*-chart. This is based on the correlation of the results of a large number of TRNSYS simulations in terms of easily calculated dimensionless variables. The *f*-chart can be used to model active liquid and air systems and active systems used to provide only hot water. These are explained in this chapter together with a modification to use the method for estimating the performance of thermosiphon solar water heating systems, which constitute the great majority of solar water heating systems. This is followed by simple models of solar energy systems suitable for analysis of the systems on hourly basis. In recent years, because of the increase of computational speed of personal computers, annual simulations have replaced design methods. Design methods however, are much faster; therefore they are still useful in early design studies. For more detailed results modeling and simulation is used. The software programs described briefly in this chapter include the F-chart and TRNSYS. Finally, the way to perform economic analysis of solar energy systems is presented and in particular the life cycle analysis. It is also indicated how to optimize a solar energy system using the economic analysis.

1. F-Chart Method

The *f*-chart is a simple method used for estimating the annual thermal performance of active heating systems using liquid or air as a working fluid and where the minimum temperature of energy delivery is near 20°C. The configurations of the systems which can be evaluated by the *f*-chart method are commonly found in residential applications. The *f*-chart method is used to estimate easily the fraction of the total heating load that can be supplied by the solar energy system called the solar contribution. If the energy required for a fuel-only system or the energy required to cover the load is denoted by $Q_{\rm L}$, the auxiliary energy for a solar energy system is denoted by $Q_{\rm AUX}$ and the solar energy delivered is $Q_{\rm S}$, then for a solar energy system:

$$Q_{\rm L} = Q_{\rm AUX} + Q_{\rm S} \tag{1}$$

For a month i, the fractional reduction of energy when a solar energy system is used, called the solar fraction or solar contribution, f, is given by the ratio:

$$f = \frac{Q_{\mathrm{L},i} - Q_{\mathrm{AUX},i}}{Q_{\mathrm{L},i}} \tag{2a}$$

which by the help of Eq. (1) is equal to:

$$f = \frac{Q_{\mathrm{S},i}}{Q_{\mathrm{L},i}} \tag{2b}$$

The *f*-chart method was originally developed by Klein and Beckman. The method is in fact a correlation of the output of hundreds of thermal performance simulations of solar

heating systems carried out with TRNSYS (see section 2.3) performed in such a way so the simulation conditions were varied over a specific range of parameters corresponding to practical system designs. These ranges are shown in Table 1. In the *f*-chart method, the primary design variable is the collector area. Secondary variables can be the storage capacity, collector type, thermal load, collector heat exchanger size and fluid flow rate. The resulting correlations give the fraction of the monthly load supplied by solar energy, *f*, as a function of two dimensionless parameters, namely:

- The ratio of collector losses to heating load, and
- The ratio of absorbed solar radiation to heating load.

Heating load includes both space heating and hot water loads. The *f*-chart method has been developed for three standard system configurations; liquid and air systems for space and hot water heating and systems for service hot water only. All these systems are active which means that they use a pump or fan for the fluid circulation and the method considers that the storage tank of the solar energy system is fully mixed. Later on the method was modified to be used for thermosiphon systems where the storage tank is stratified and the flow in the collector occurs by natural means.

Parameter	Range
Transmittance-absorptance product at normal incidence - $(\tau \alpha)_n$	0.6-0.9
Product of heat removal factor with heat exchanger and collector	
area - $F'_R A_c$	$5-120 \text{ m}^2$
Collector heat loss coefficient - U _L	$2.1-8.3 \text{ W/m}^2-^{\circ}\text{C}$
Collector slope from horizontal - β	30-90°
Overall energy loss coefficient and area product for a building	
used in a degree day space heating load - (UA) _h	83.3-666.6 W/°C

Table 1. Range of design variables used in developing *f*-charts for liquid and air systems

The solar energy delivered $Q_{\rm S}$, can be obtained by integrating over a time period Δt the useful energy collected by a solar collector system. Thus using the definition of the solar fraction, f, from Eq. (2b), we get:

$$f = \frac{Q_{\rm S}}{Q_{\rm L}} = \frac{1}{Q_{\rm L}} \int_{\Delta t} Q_{\rm u}^+ dt \tag{3}$$

where $Q_{\rm L}$ = total heating load during the integration period (MJ)

The useful energy collected by the solar energy system Q_u , is obtained by the following relation:

$$Q_{\rm u} = A_{\rm c} F_{\rm R} \Big[G_{\rm t}(\tau \alpha) - U_{\rm L} \big(T_{\rm i} - T_{\rm a} \big) \Big], \tag{4}$$

where $F_{\rm R}$ is the heat removal factor, which represents the ratio of the actual useful energy gain that would result if the collector-absorbing surface had been at the local fluid temperature, given by:

$$F_{\rm R} = \frac{\dot{m}c_{\rm p}}{A_{\rm c}U_{\rm L}} \left(1 - \exp\left[-\frac{U_{\rm L}F'A_{\rm c}}{\dot{m}c_{\rm p}} \right] \right).$$
(5)

In Eq. (4) by replacing F_R by F'_R , to account for the presence of a heat exchanger in the solar collection network as is explained subsequently, and G_t , the total incident radiation (beam and diffuse), by H_t , the total insolation over a day, Eq. (3) can be written as:

$$f = \frac{A_{\rm c} F_{\rm R}}{Q_{\rm L}} \int_{\Delta t} \left[H_{\rm t}(\tau \alpha) - U_{\rm L}(T_{\rm s} - T_{\rm a}) \right] dt \,.$$
(6)

The last term of Eq. (6) can be multiplied and divided by the term $(T_{ref} - T_a)$, where T_{ref} is a reference temperature chosen to be 100°C, so Eq. (6) can be written as:

$$f = \frac{A_{\rm c} F_{\rm R}'}{Q_{\rm L}} \int_{\Delta t} \left[H_{\rm t}(\tau \alpha) - U_{\rm L} \left(T_{\rm ref} - T_{\rm a} \right) \frac{\left(T_{\rm s} - T_{\rm a} \right)}{\left(T_{\rm ref} - T_{\rm a} \right)} \right] dt \,. \tag{7}$$

In any solar energy system the resulting storage tank temperature T_s , is a complicated function of H_t , Q_L and T_a , therefore Eq. (7) cannot be explicitly integrated and evaluated. It suggests however, that an empirical correlation can be found, on a monthly basis, between the solar fraction, f, and the two dimensionless groups mentioned earlier as follows:

$$X = \frac{A_{\rm c}F_{\rm R}'U_{\rm L}}{Q_{\rm L}}\int_{\Delta t} \left(T_{\rm ref} - \overline{T_{\rm a}}\right)dt = \frac{A_{\rm c}F_{\rm R}'U_{\rm L}}{Q_{\rm L}} \left(T_{\rm ref} - \overline{T_{\rm a}}\right)\Delta t \tag{8}$$

$$Y = \frac{A_{\rm c}F_{\rm R}'}{Q_{\rm L}}\int_{\Delta t} H_{\rm t}(\tau\alpha)dt = \frac{A_{\rm c}F_{\rm R}'}{Q_{\rm L}}(\overline{\tau\alpha})\overline{H}_{\rm t}N, \qquad (9)$$

where $Q_{\rm L}$ = monthly heating load or demand (MJ)

N = number of days in a month

 \overline{T}_{a} = monthly average ambient temperature (°C)

 \overline{H}_{t} = monthly average daily total radiation on the tilted collector surface (MJ/m²)

 $(\tau \alpha)$ = monthly average value of $(\tau \alpha)$, which is equal to the monthly average value of absorbed over incident solar radiation, given by:

$$(\overline{\tau\alpha}) = \frac{\overline{S}}{\overline{H}_{t}}$$
(10)

The dimensionless parameters X and Y, given by Eqs. (8) and (9) are usually rearranged to use the factors $F_{\rm R}U_{\rm L}$ and $F_{\rm R}(\tau\alpha)_{\rm n}$ which are readily available from standard collector tests and therefore they become:

$$X = F_{\rm R} U_{\rm L} \frac{F_{\rm R}'}{F_{\rm R}} (T_{\rm ref} - \overline{T}_{\rm a}) \Delta t \frac{A_{\rm c}}{Q_{\rm L}}$$
(11)

$$Y = F_{\rm R}(\tau\alpha)_{\rm n} \frac{F_{\rm R}'}{F_{\rm R}} \left[\frac{(\overline{\tau\alpha})}{(\tau\alpha)_{\rm n}} \right] \overline{H}_{\rm t} N \frac{A_{\rm c}}{Q_{\rm L}}$$
(12)

The physical significance of the two dimensionless parameters X and Y is that the parameter X represents the ratio of the reference collector total energy loss to total heating load or demand (Q_L) during the period Δt , whereas the parameter Y represents the ratio of the total absorbed solar energy to the total heating load or demand (Q_L) during the same period.

The ratio F'_R / F_R in Eqs. (11) and (12) is used to correct the collector performance to account for the presence of a heat exchanger which causes the collector side of the system to operate at higher temperature than a similar system without a heat exchanger and is given by:

$$\frac{F_{\rm R}'}{F_{\rm R}} = \left[1 + \frac{A_{\rm c}F_{\rm R}U_{\rm L}}{\left(\dot{m}c_{\rm p}\right)_{\rm c}} \left(\frac{\left(\dot{m}c_{\rm p}\right)_{\rm c}}{\varepsilon\left(\dot{m}c_{\rm p}\right)_{\rm min}} - 1\right)\right]^{-1}$$
(13)

where $(\dot{m}c_p)_{min}$ = smaller of the fluid capacitance rates of the collector and tank sides of the heat exchanger (W/°C) ε = heat exchanger effectiveness (-).

For a given collector orientation, the value of the factor $(\overline{\tau\alpha})/(\tau\alpha)_n$ in Eq. (12) varies slightly from month to month. For collectors tilted and facing the equator with a slope equal to latitude plus 15°, Klein found that the factor is equal to 0.96 for one-cover collector and 0.94 for two-covers collectors for the whole heating season (winter months). Using Eq. (10) for the relation of $(\overline{\tau\alpha})$:

$$\frac{\left(\overline{\tau\alpha}\right)}{\left(\tau\alpha\right)_{\rm n}} = \frac{\overline{S}}{\overline{H}_{\rm t}\left(\tau\alpha\right)_{\rm n}}.$$
(14)

If the isotropic model is used, \overline{S} is given by:

$$\overline{S} = \overline{H}_{B}\overline{R}_{B}(\overline{\tau\alpha})_{B} + \overline{H}_{D}(\overline{\tau\alpha})_{D}\left(\frac{1+\cos(\beta)}{2}\right) + \overline{H}\rho_{G}(\overline{\tau\alpha})_{G}\left(\frac{1-\cos(\beta)}{2}\right).$$
(15)

Therefore, by substituting in Eq. (14) gives:

$$\frac{\left(\overline{\tau\alpha}\right)}{\left(\tau\alpha\right)_{n}} = \frac{\overline{H}_{B}\overline{R}_{B}}{\overline{H}_{t}} \frac{\left(\overline{\tau\alpha}\right)_{B}}{\left(\tau\alpha\right)_{n}} + \frac{\overline{H}_{D}}{\overline{H}_{t}} \frac{\left(\overline{\tau\alpha}\right)_{D}}{\left(\tau\alpha\right)_{n}} \left(\frac{1+\cos(\beta)}{2}\right) + \frac{\overline{H}\rho_{G}}{\overline{H}_{t}} \frac{\left(\overline{\tau\alpha}\right)_{G}}{\left(\tau\alpha\right)_{n}} \left(\frac{1-\cos(\beta)}{2}\right).$$
(16)

The term $\overline{R}_{\rm B}$ is the ratio of the monthly average beam radiation on a tilted surface to that on a horizontal surface, called the monthly mean beam radiation tilt factor. This is a complicated function of the atmospheric transmittance, but it can be estimated by the ratio of extraterrestrial radiation on the tilted surface to that on a horizontal surface for the month. For surfaces facing directly towards the equator this is given by:

$$R'_{\rm B} = \frac{\cos(L-\beta)\cos(\delta)\sin(h'_{\rm ss}) + (\pi/180)h'_{\rm ss}\sin(L-\beta)\sin(\delta)}{\cos(L)\cos(\delta)\sin(h_{\rm ss}) + (\pi/180)h_{\rm ss}\sin(L)\sin(\delta)}$$
(17a)

where h'_{ss} is sunset hour angle on the tilted surface (degrees) given by:

$$h'_{\rm ss} = \min\{h_{\rm ss}, \cos^{-1}\left[-\tan\left(L-\beta\right)\tan\left(\delta\right)\right]\}.$$
(17b)

For the southern hemisphere the term $(L-\beta)$ of Eqs. (17a) and (17b) need to change to $(L+\beta)$.

In Eq. (16) the $(\tau \alpha)/(\tau \alpha)_n$ ratios can also be obtained from graphs given in the original paper of Klein or from the following curve fit equations:

For single cover:

$$\frac{(\overline{\tau\alpha})}{(\tau\alpha)_{\rm n}} = -8.7x10^{-8}\theta^4 + 1.03x10^{-5}\theta^3 - 0.0004762\theta^2 + 0.00851\theta + 0.94967.$$
(18a)

For 2 covers:

$$\frac{(\overline{\tau\alpha})}{(\tau\alpha)_{\rm n}} = -5.05x10^{-8}\theta^4 + 3.578x10^{-6}\theta^3 - 8.777x10^{-5}\theta^2 - 1.836x10^{-6}\theta + 1.0042, \quad (18b)$$

where θ is the incidence angle in degrees.

Equations (18a) and (18b) are for one and two sheet of collector glass covers respectively with refractive index equal to 1.526. The various components

of $(\tau \alpha)/(\tau \alpha)_n$ in Eq. (16), i.e., the beam, diffuse and ground reflected components are obtained from Eqs. (18a) and (18b) by using the appropriate mean incidence angle. For the beam component the effective angle of incidence $\overline{\theta}_b$, also called mean incidence angle for beam radiation, can be obtained from the original figures of Klein and for surfaces facing directly towards the equator it can be approximated as the incidence angle at 2.5 hours from solar noon on the average day of the month shown in Table 2 together with other useful parameters like the day number, the declination at that day, δ , and β_m , the monthly optimal collector tilt angle.

Month	Day	Day No.	δ (deg.)	β_m (deg.)	
January	17	17	-20.92	L+29	
February	16	47	-12.95	L+18	
March	16	75	-2.42	L+3	
April	15	105	9.41	L-10	
May	15	135	18.79	L-22	
June	11	162	23.09	L-25	
July	17	198	21.18	L-24	
August	16	228	13.45	L-10	
September	15	258	2.22	L-2	
October	15	288	-9.60	L+10	
November	14	318	-18.91	L+23	
December	10	344	-23.05	L+30	

 Table 2. Recommended average day for each month, declination for that day and monthly optimal tilt angle

For the diffuse and ground reflected components of radiation at the effective incidence angles at β , the mean incidence angles can be obtained from:

$$\theta_{\rm e,D} = 59.68 - 0.1388\beta + 0.001497\beta^2 \tag{19}$$

$$\theta_{\rm e,G} = 90 - 0.5788\beta + 0.002693\beta^2 \tag{20}$$

Generally, *f*-chart is used to estimate the monthly solar fraction, f_i . The energy contribution for the month, *i*, is the product of f_i and monthly load (heating and hot water), $Q_{L,i}$, i.e., $f_i Q_{L,i}$. The fraction on an annual basis of the annual load supplied by the solar energy system, *F*, is the sum of the monthly energy contributions divided by the annual load, given by:

$$F = \frac{\sum f_i Q_{L,i}}{\sum Q_{L,i}} \tag{21}$$

It should be noted that although the method primarily depends on values obtained from charts, hence its name, here only the main f-charts and equations are given. Other

complimentary charts can be found from the original references or from books. However, the interested reader can easily create the appropriate charts using any spreadsheet program or the spreadsheet program can be used to obtain the required values easily and more accurately without having to perform a large number of hand calculations.

It should be pointed out again that the method can be used to simulate standard solar water and air systems configurations and solar energy systems used only for hot water production. These are examined separately in the following sections.

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Biographical Sketch

Dr. Soteris Kalogirou is a Senior Lecturer at the Department of Mechanical Engineering and Materials Sciences and Engineering of the Cyprus University of Technology, Limassol, Cyprus. He received his HTI Degree in Mechanical Engineering in 1982, his M.Phil. in Mechanical Engineering from the Polytechnic of Wales in 1991 and his Ph.D. in Mechanical Engineering from the University of Glamorgan in 1995. In June 2011 he received from the University of Glamorgan the title of D.Sc. For more than 25 years, he is actively involved in research in the area of solar energy and particularly in flat plate and concentrating collectors, solar water heating, solar steam generating systems, desalination and absorption cooling. Additionally, since 1995 he is involved in a pioneering research dealing with the use of artificial intelligence methods, like artificial neural networks, genetic algorithms and fuzzy logic, for the modeling and performance prediction of energy and solar energy systems. He has 34 books and book contributions and published 233 papers; 99 in international scientific journals and 134 in refereed conference proceedings. Until now, he received more than 2750 citations on this work. He is Executive Editor of *Energy*, Associate Editor of *Renewable Energy* and Editorial Board Member of another eleven journals. He is the editor of the book *Artificial Intelligence in Energy and Renewable Energy Systems*, published by Nova Science Inc., co-editor of the book *Soft Computing in Green and Renewable Energy Systems*, published

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